

# Climate Science 101

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MIT IAP 2017



# Purpose of these lectures

- Develop a broad understanding of Earth's climate system
- Understand how Earth's climate responds to changes that are natural and human-induced

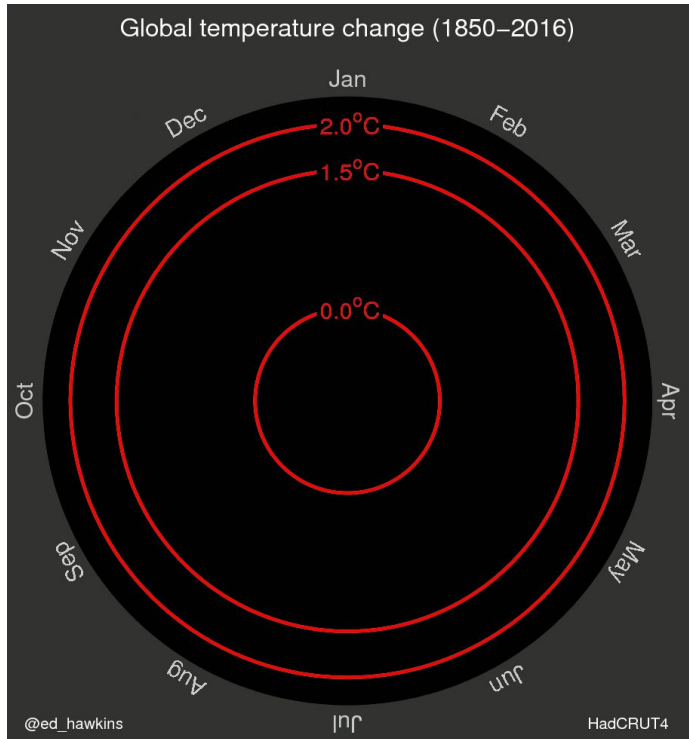
# What these lectures are not

- Not about the politics or policies surrounding the issue of climate change

# Today's topics

- History of climate science
- Structure and composition of Earth's atmosphere
- Earth's energy budget
- The greenhouse effect
- Variability in the climate system
- Emissions of greenhouse gases and their long-lived persistence

# The world is warming. Why?



Global temperature change since 1850 to present day.

Since 1850, the global mean temperature has increased by around 0.8 degrees Celsius.

# Definition of “Climate”

## Popular definitions:

- Climate is the average weather
- Climate is what you expect, weather is what you get

## Scientific definition:

- **Climate is the statistics of weather**
  - Not just the mean weather, but the statistics of its variability
  - Aggregation over time scales of more than one year so that seasonal cycle is not considered

## Examples climate variability:

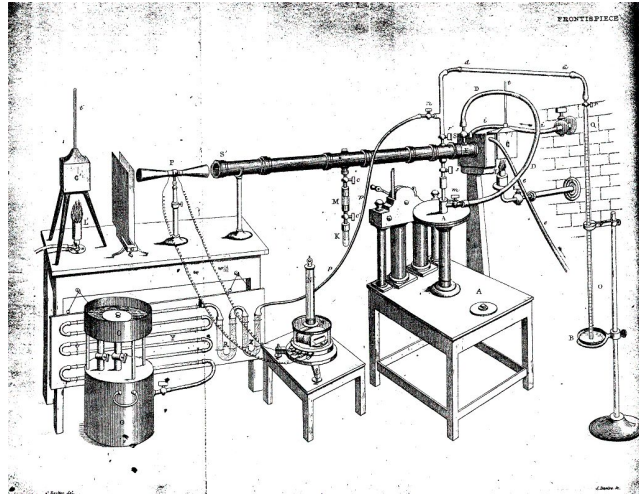
- El Nino/La Nina (2-4y), “Little ice age” (100y), Glacial cycles (20,000y)

# History of climate science

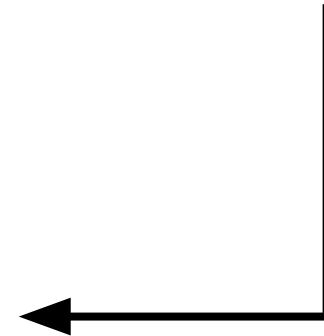


Jean Baptiste  
Joseph Fourier  
(1768-1830)

## The Greenhouse Effect



John Tyndall  
(1820-1893)



# History of climate science



Svante Arrhenius  
1859-1927

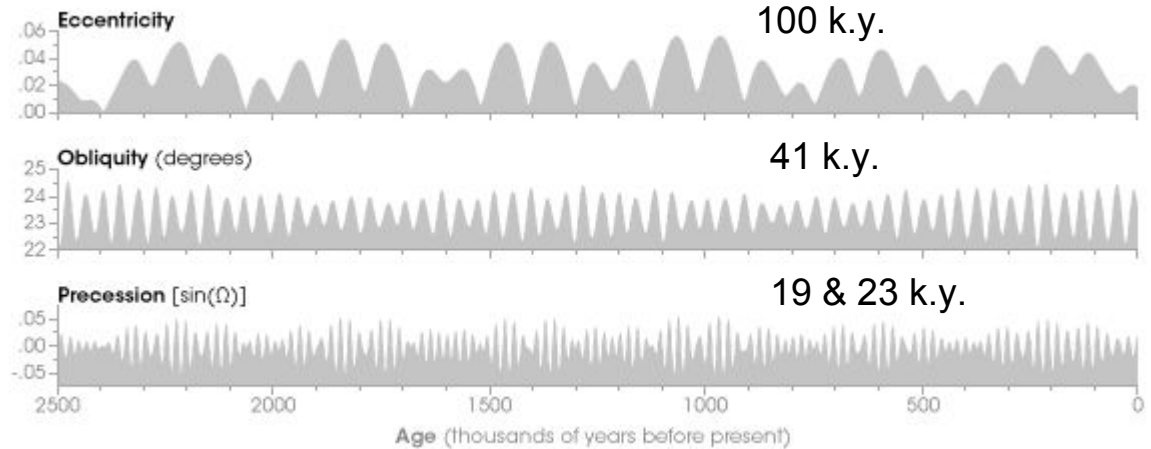
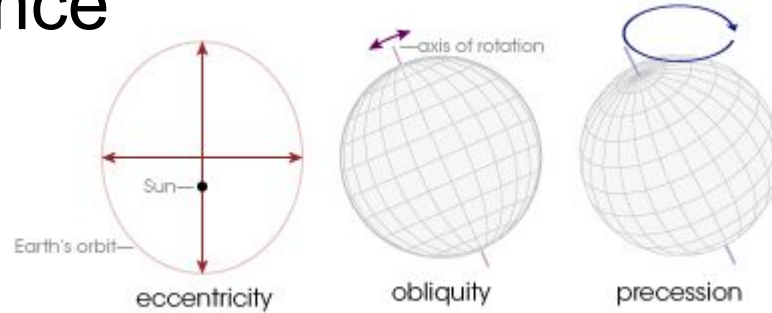
*"If the quantity of carbonic acid [CO<sub>2</sub>] in the air should sink to one-half its present percentage, the temperature would fall by about 4°; a diminution to one-quarter would reduce the temperature by 8°. **On the other hand, any doubling of the percentage of carbon dioxide in the air would raise the temperature of the earth's surface by 4°; and if the carbon dioxide were increased fourfold, the temperature would rise by 8°.**"*

Världarnas utveckling, 1906. p.53

# History of climate science

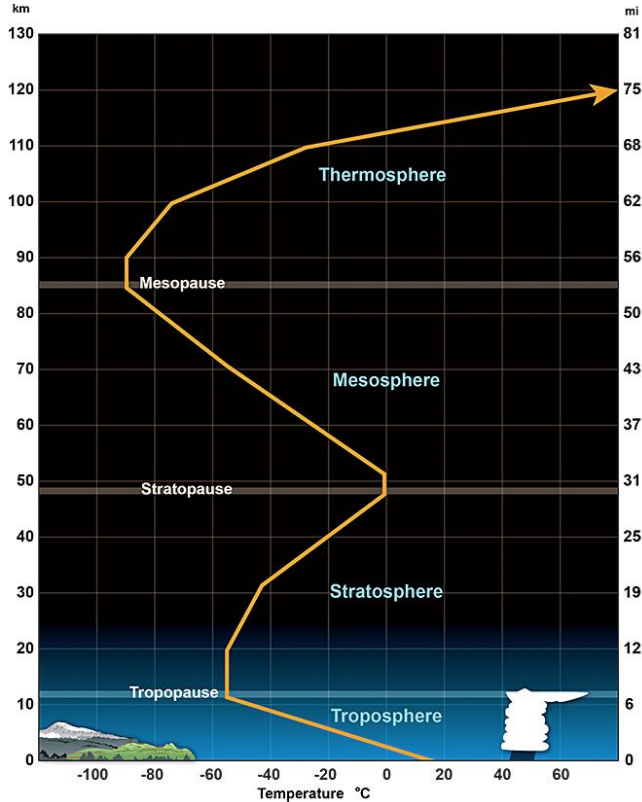


Milutin Milankovic  
1879-1958





# Structure of the atmosphere

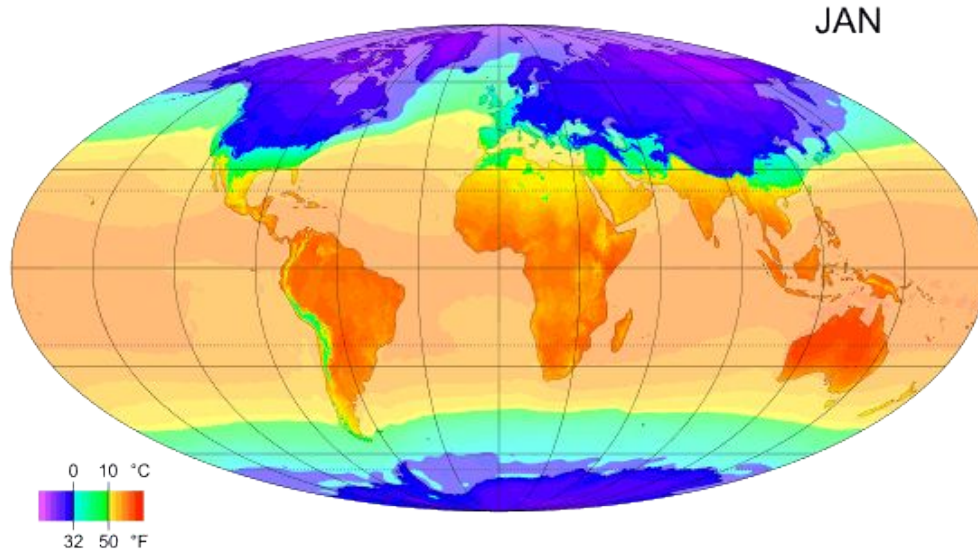


← Thermosphere: hot from solar charged particle interactions

← Middle atmosphere: stratosphere and mesosphere

← Troposphere: 'weather' layer

# Temperature climatology

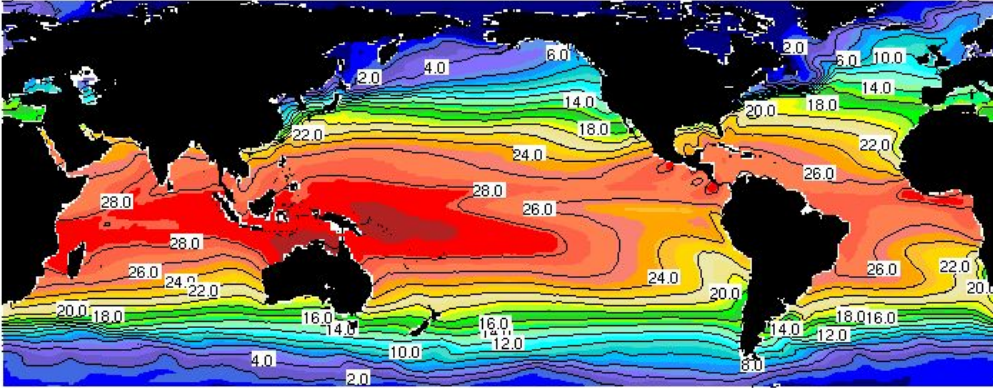


Mean surface temperatures:

- Warm temperature belt in tropics migrates North and South with the seasons.
- Most extreme variations are in high latitudes of Northern hemisphere

# Temperature climatology

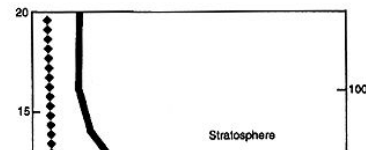
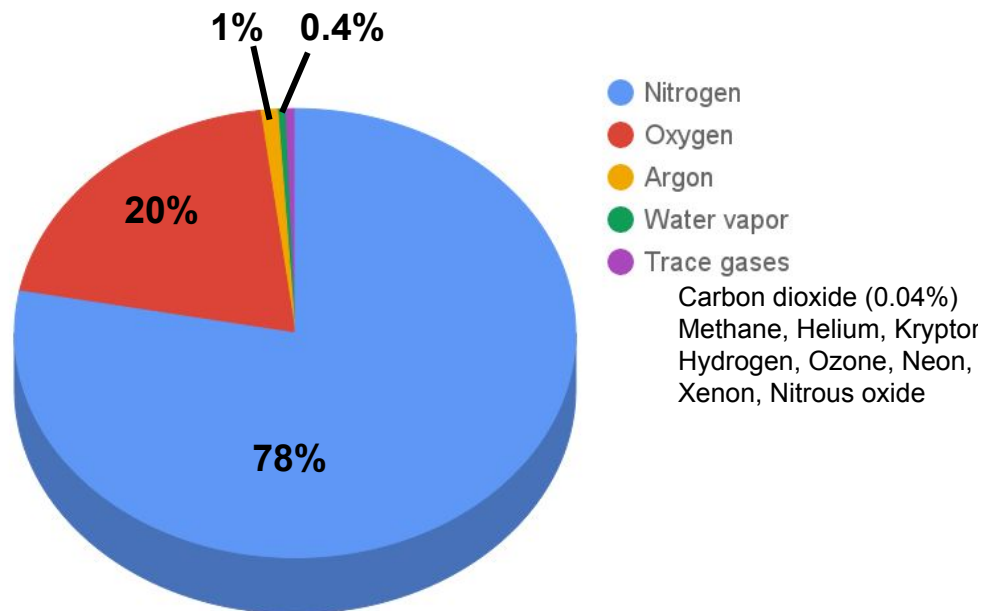
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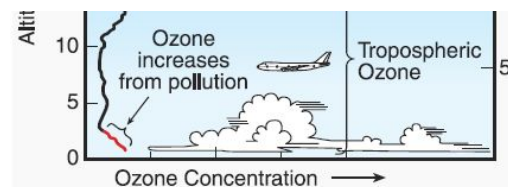
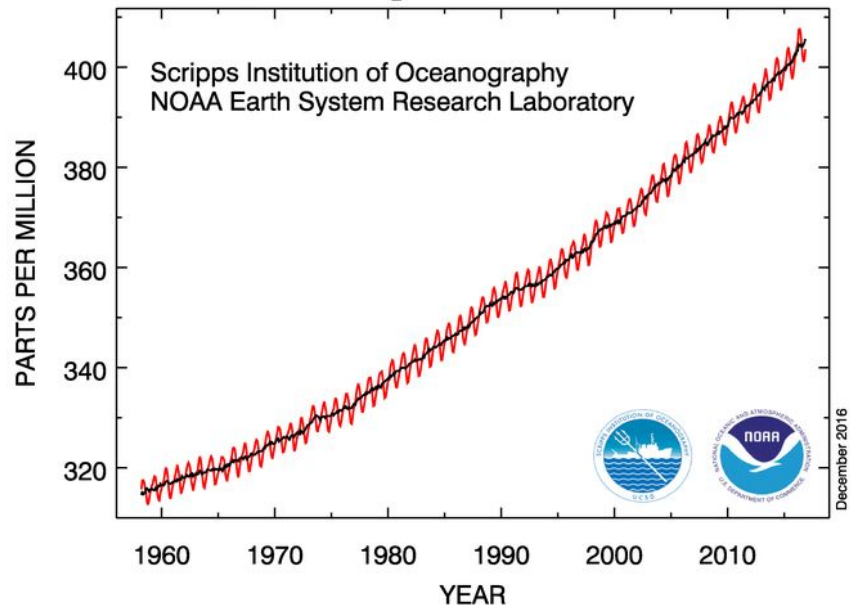
Sea surface temperatures:

- Highest temperatures around equator, but not zonally symmetric
- Many interesting features/currents you can see

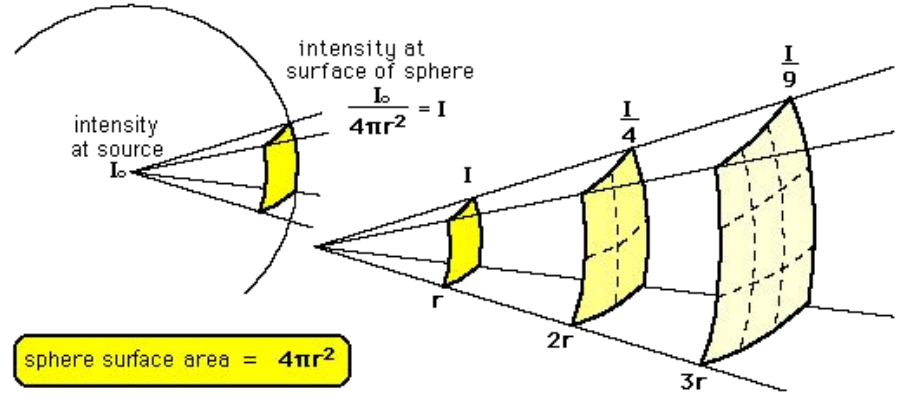
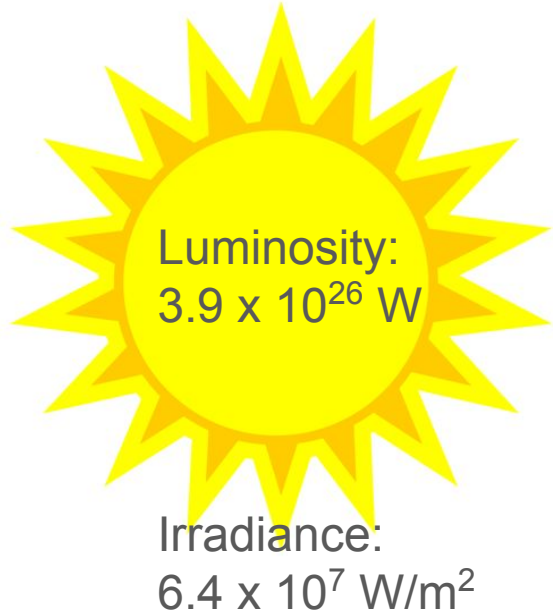
# Atmospheric composition



Atmospheric CO<sub>2</sub> at Mauna Loa Observatory



# Earth energy balance



$1.5 \times 10^{11} \text{ m}$

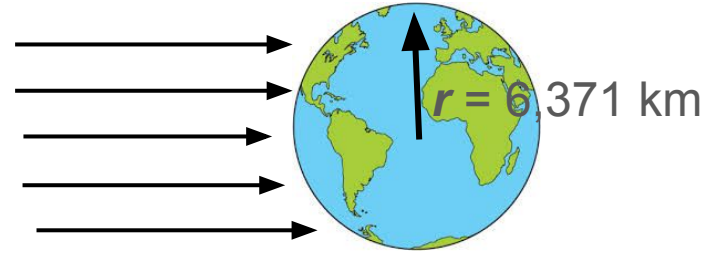


Irradiance:  
 $1370 \text{ W/m}^2$



Solar constant  $S_0$

# Earth energy balance



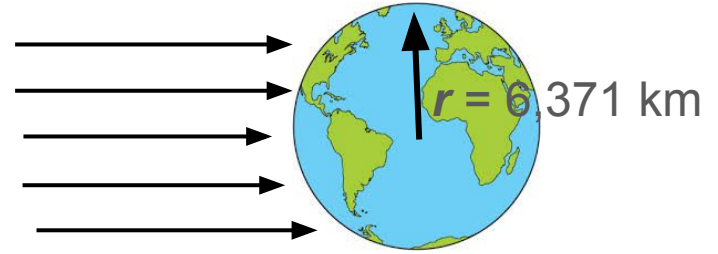
Total absorbed solar radiation:  $S_0(1-a_p)\pi r^2$        $a_p \simeq 0.30$

Absorption per unit area:  $\frac{S_0}{4}(1-a_p)$

Planetary albedo is the fraction of incident radiation reflected to space

# Earth energy balance

Case of no atmosphere:



Can estimate Earth's surface temperature using Stefan Boltzmann law:

$$F = \sigma T^4 \quad \sigma = 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^{-4}$$

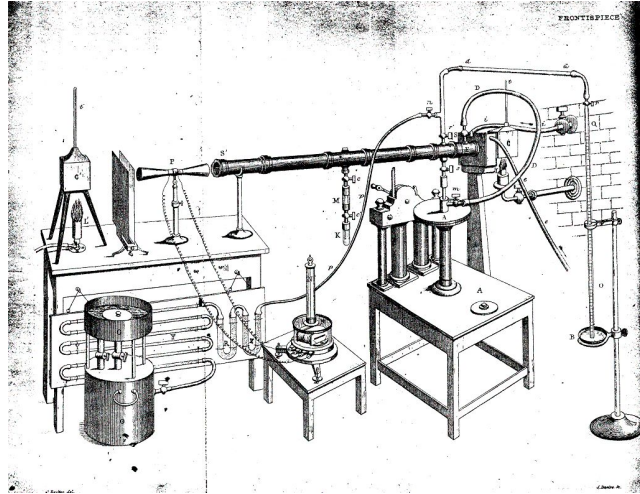
$$\sigma T_e^4 = \frac{S_0}{4} (1 - a_p)$$

The effective emission temperature of Earth is 255 K (-18°C) if we had no atmosphere. Observed average surface temperature is 15°C. **Way too cold!!!**

# Greenhouse effect

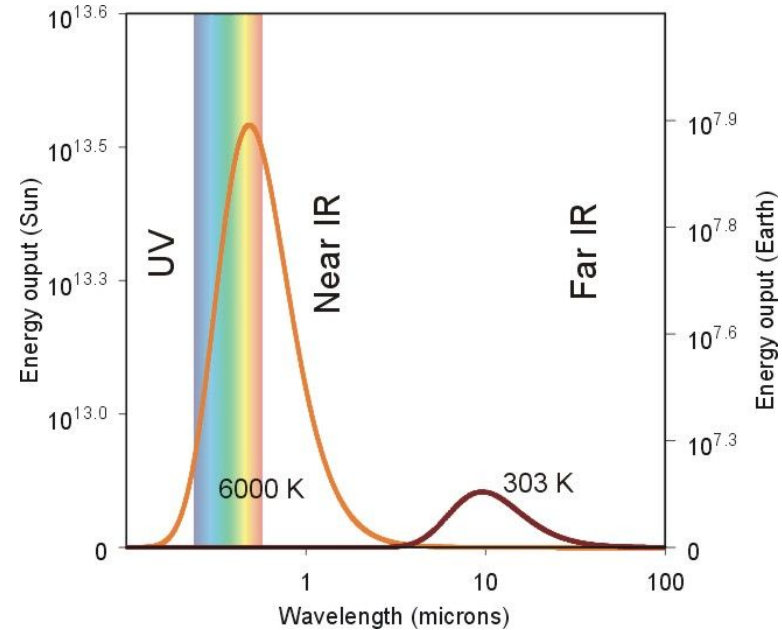


John Tyndall  
(1820-1893)



Measured infrared (IR) absorption of atmospheric gases

## Wien's law



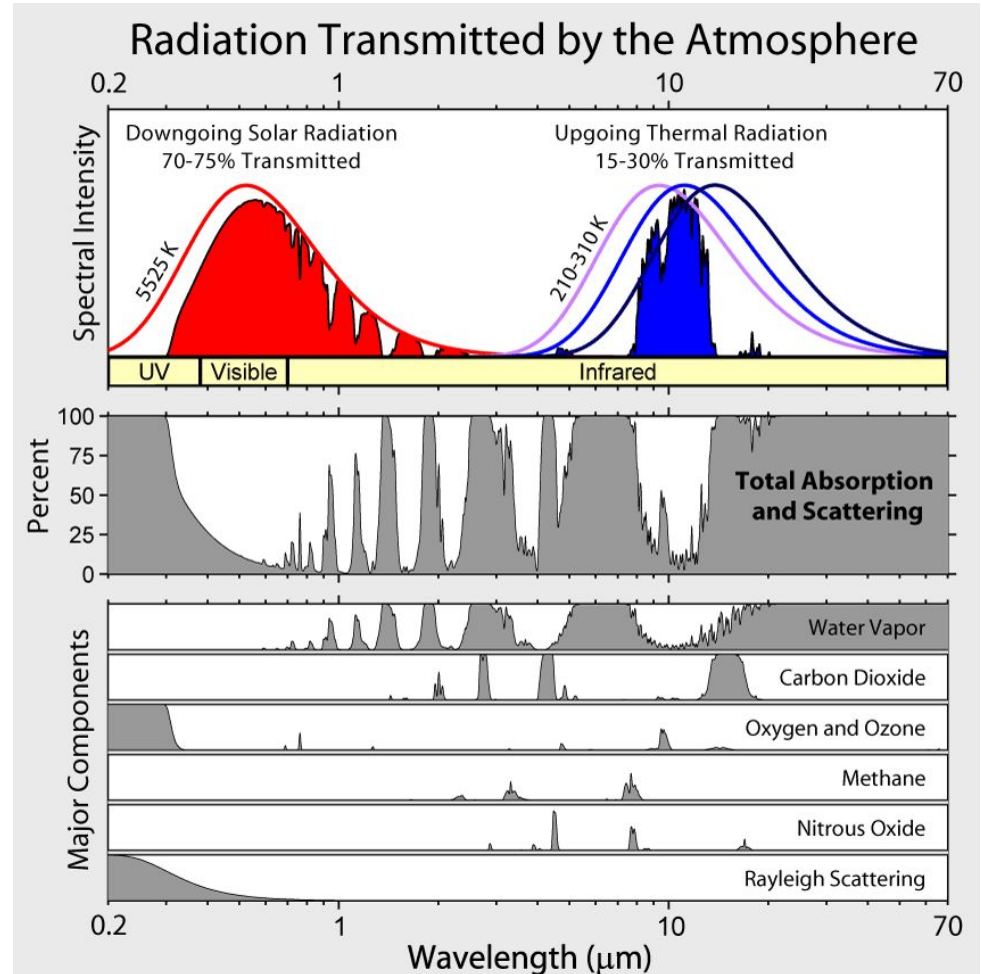


# Greenhouse effect

Solid lines are the Planck blackbody emission spectra, as function of temperature.

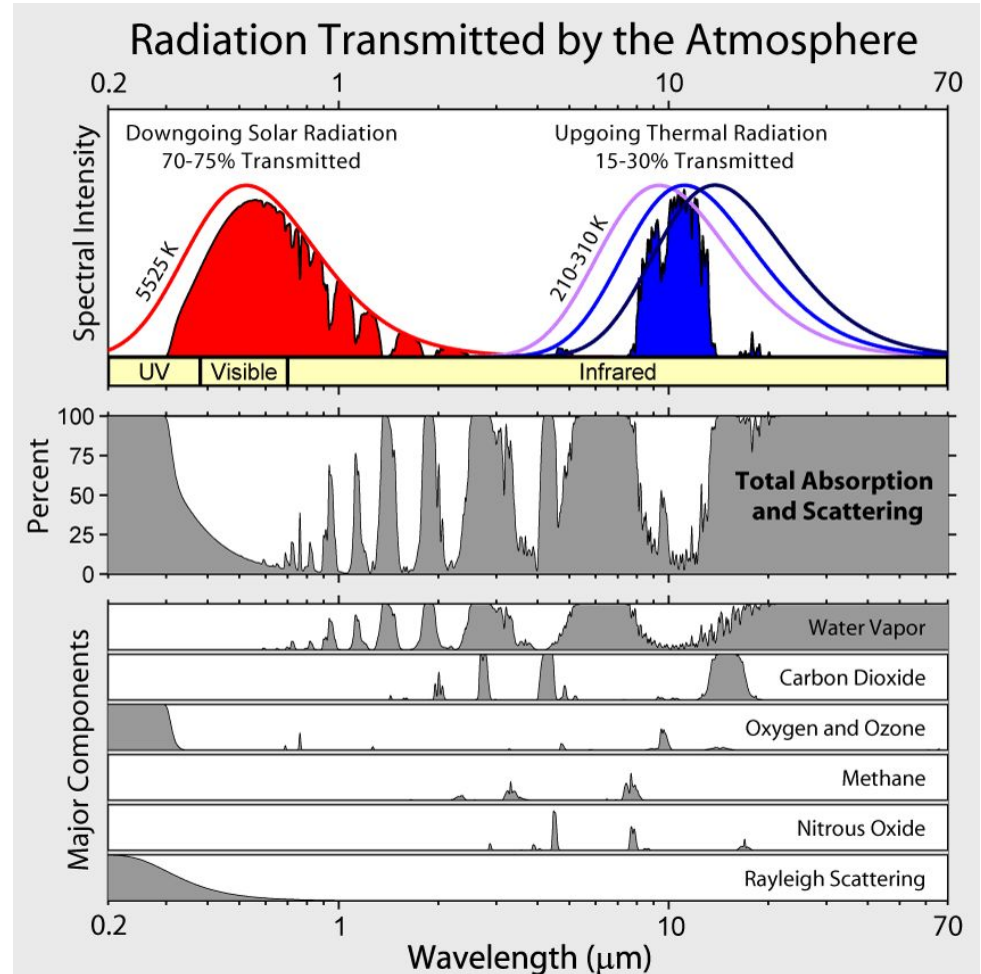
The solid filled parts show the spectra of the **incoming** solar radiation transmitted through the atmosphere to the surface, and the **outgoing** terrestrial radiation transmitted through to the top of the atmosphere.

The difference between the two shows the wavelengths at which the solar/terrestrial radiation is absorbed, and by what gas/process.



# Greenhouse effect

- Only 15-30% of terrestrial IR radiation escapes to space, with almost all the far IR being absorbed by the atmosphere.
- Water vapor is strongest absorber of IR heat like Tyndall concluded.
- Carbon dioxide bands do not completely overlap with water vapor.



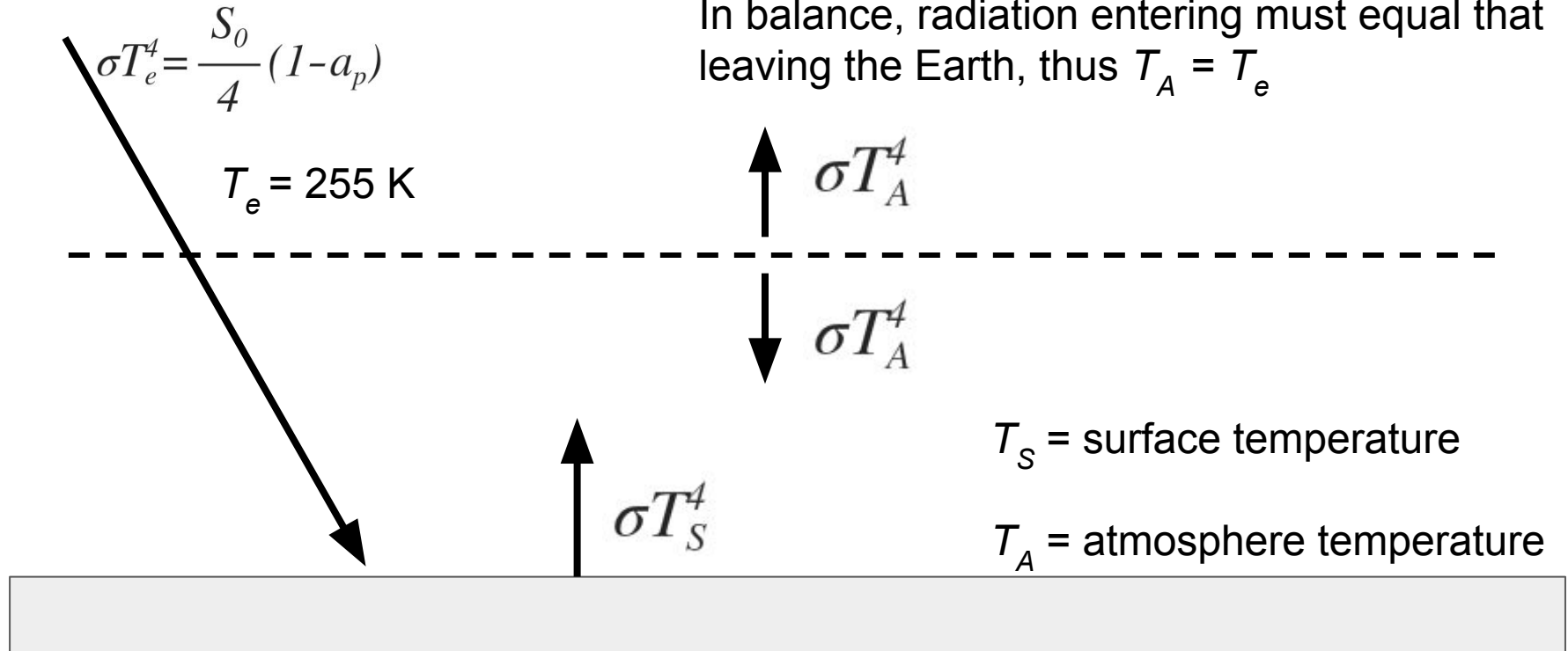
# Simple model with the greenhouse effect

Assumptions:

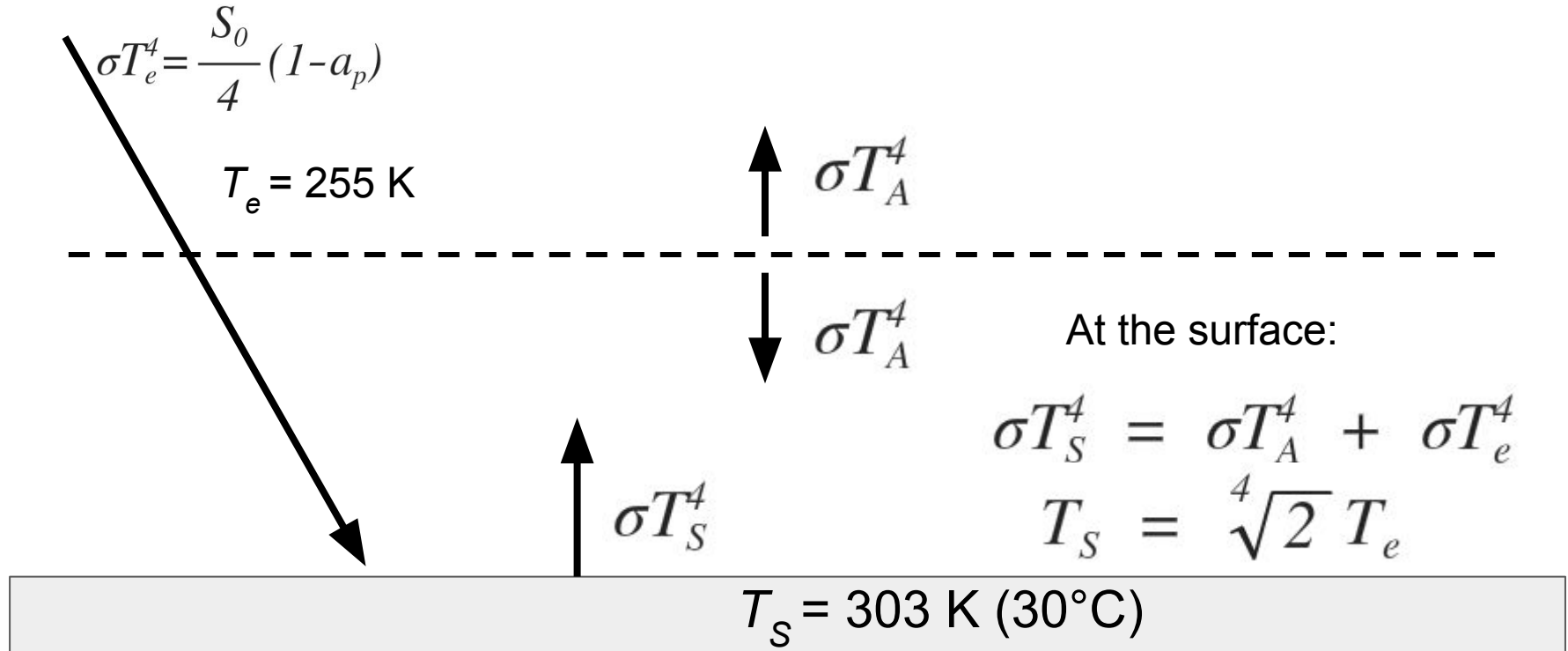
- Atmosphere is completely transparent to solar radiation
- Atmosphere is opaque to infrared radiation
- Infrared emission is from surface and the atmospheric layer
- Atmosphere is a single slab

# Simple model with the greenhouse effect

In balance, radiation entering must equal that leaving the Earth, thus  $T_A = T_e$



# Simple model with the greenhouse effect



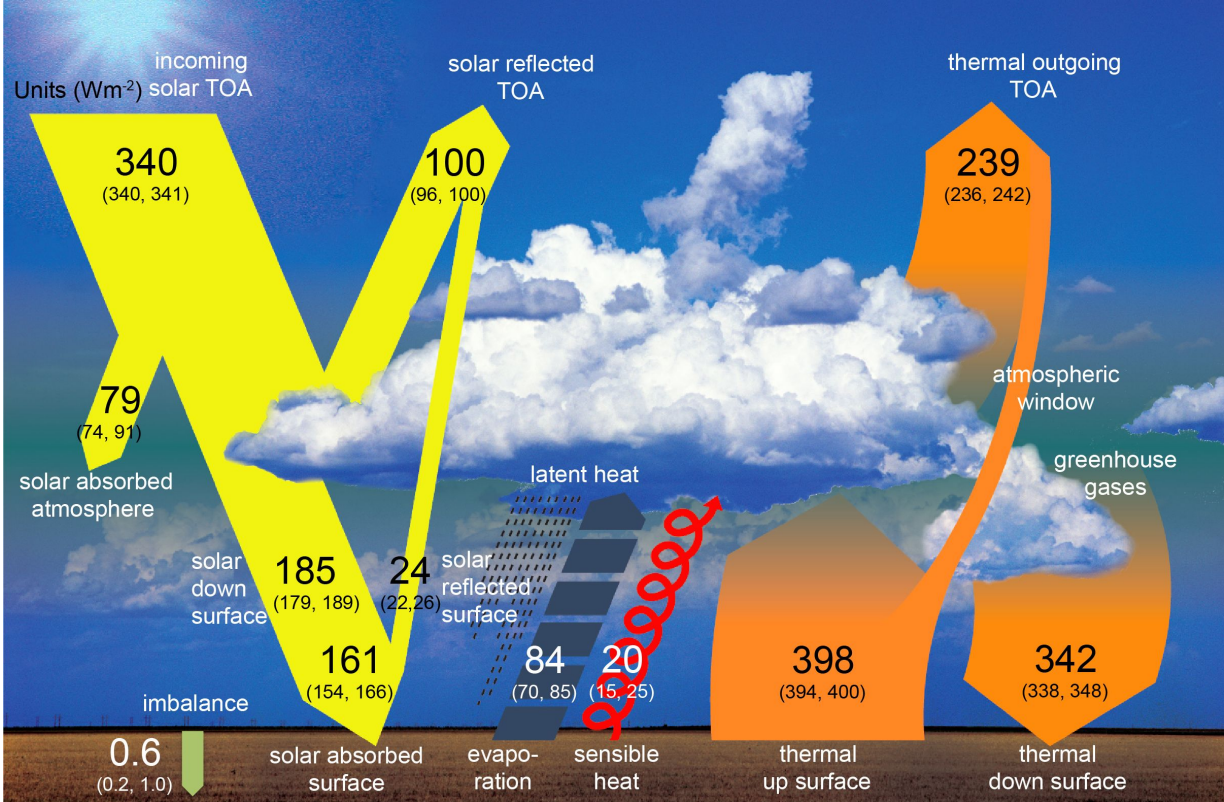
# Simple model with the greenhouse effect

- Due to greenhouse effect, the surface must warm enough to balance both incoming sunlight and the radiation from the atmosphere back to the surface

## Limitations:

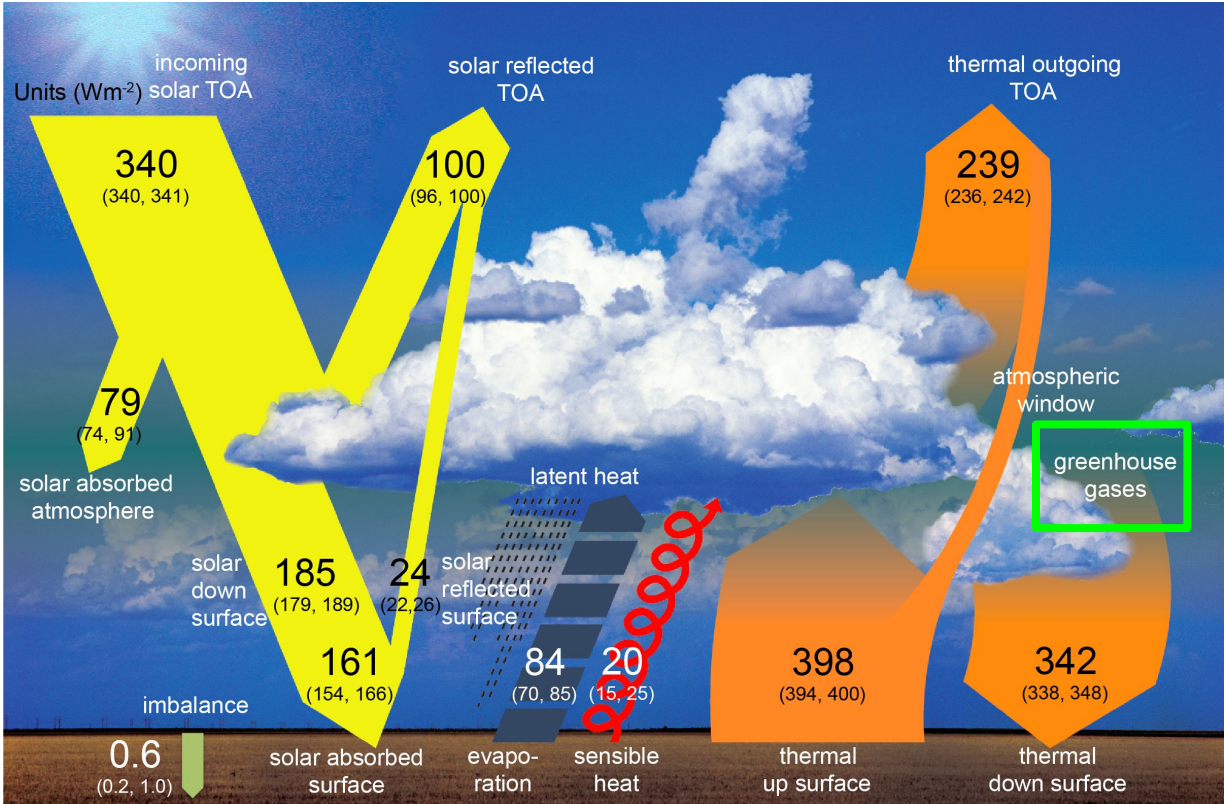
- Atmosphere is not completely opaque to infrared radiation, consider 'leaky' greenhouse, and more than one layer
- Heat transported by convection as well as radiation

# Earth's radiative budget



Global mean energy budget

# Earth's radiative budget



Increasing greenhouse gases, increases IR emission back to the surface, which must adjust, i.e. by warming up.



# Climate variability

## Causes of natural climate variability

- 'External forcing' of the climate system:
  - Glacial periods caused by changes in Earth's orbit ~10,000 - 100,000 years
  - Large volcanic eruptions
  - Solar variability
- 'Internal' climate variability (chaos due to nonlinear interactions of a complex system)
  - El Niño, North Atlantic Oscillation

# Volcanic eruptions impact climate

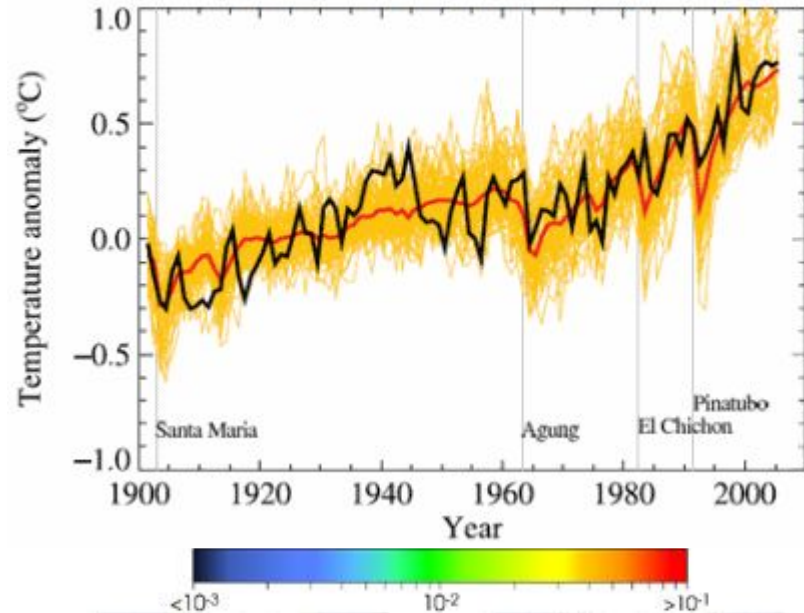
Mount Pinatubo, Philippines 1991



SAGE II 1020 nm Optical Depth

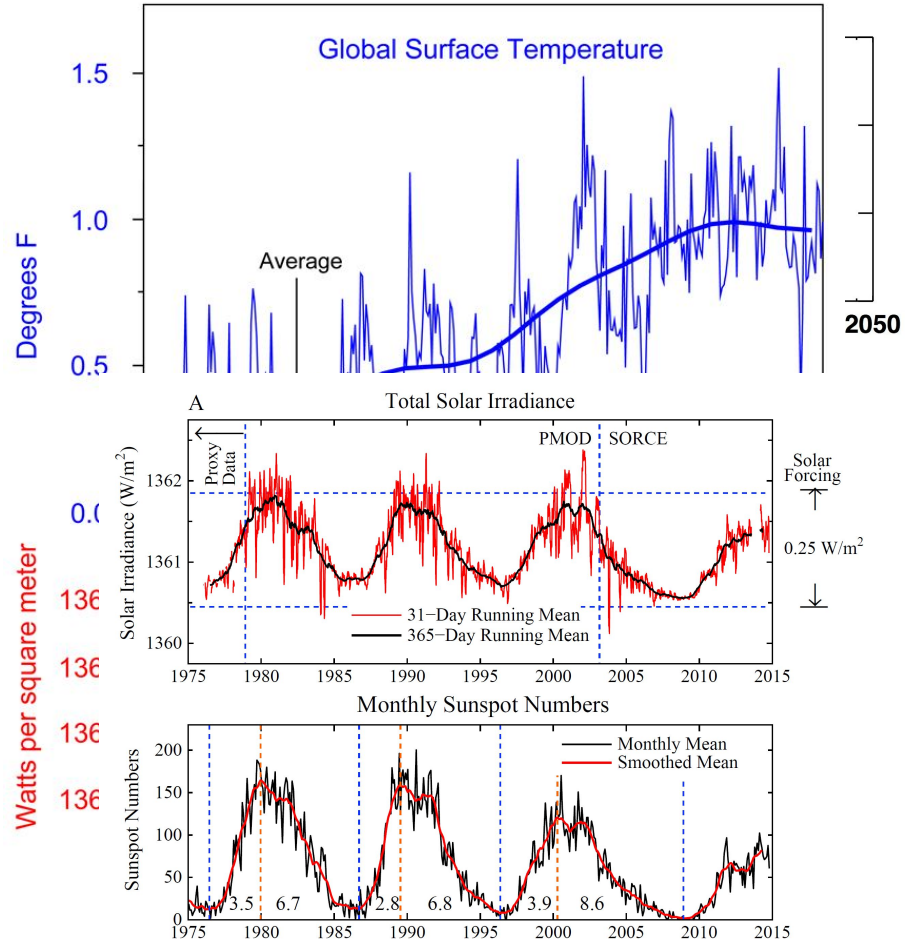
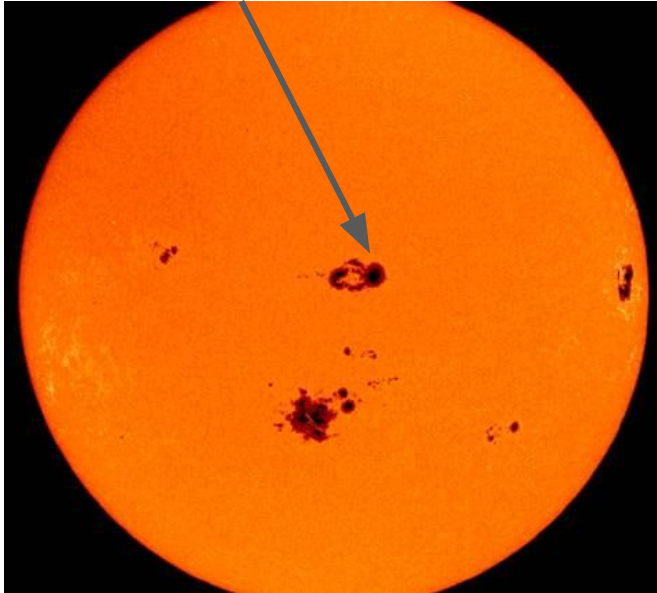


Volcanoes can have cooling effects



# Solar forcing of climate

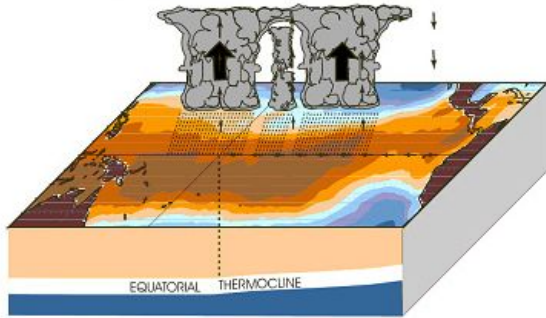
Energy output of the Sun shows slight fluctuations as the number of sunspots varies



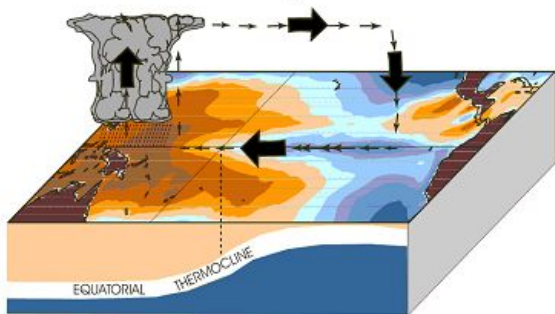
# El Niño Southern Oscillation (ENSO)

The ENSO cycle describes the fluctuations in temperature between the ocean and atmosphere in the Pacific that lead to large-scale weather pattern changes across many regions of the world.

December - February El Niño Conditions



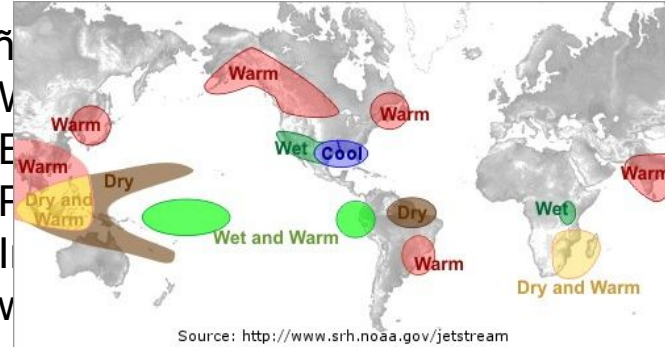
December - February La Niña Conditions



NOAA/NCEP/CPC

## El Niño

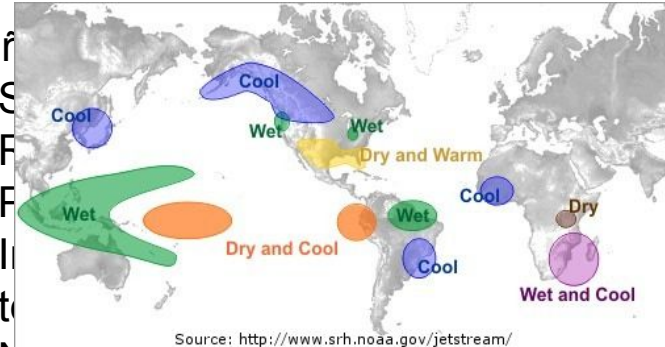
- V
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- I
- V



Source: <http://www.srh.noaa.gov/jetstream>

## La Niña

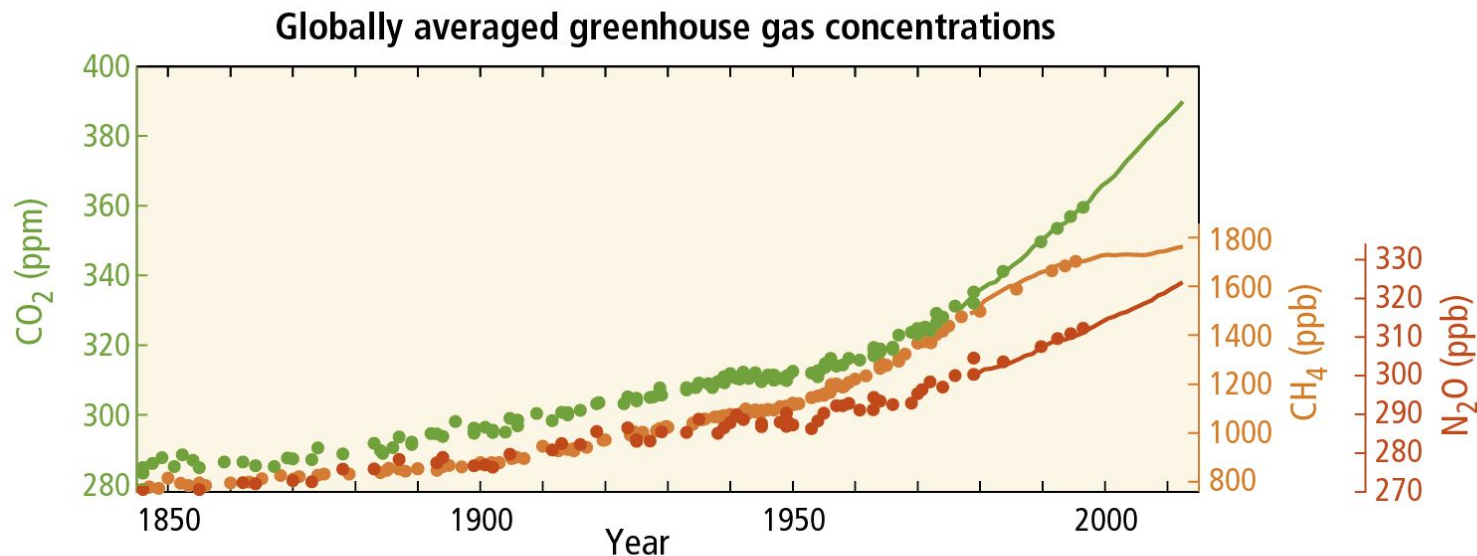
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Source: <http://www.srh.noaa.gov/jetstream/>

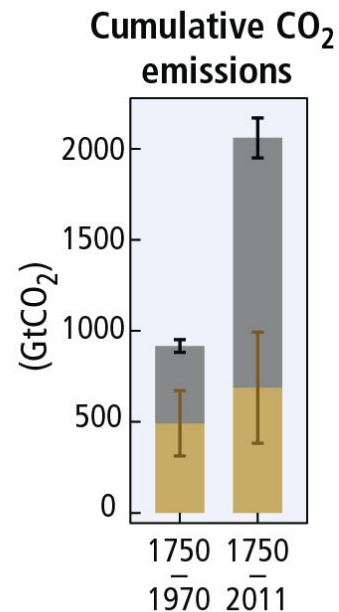
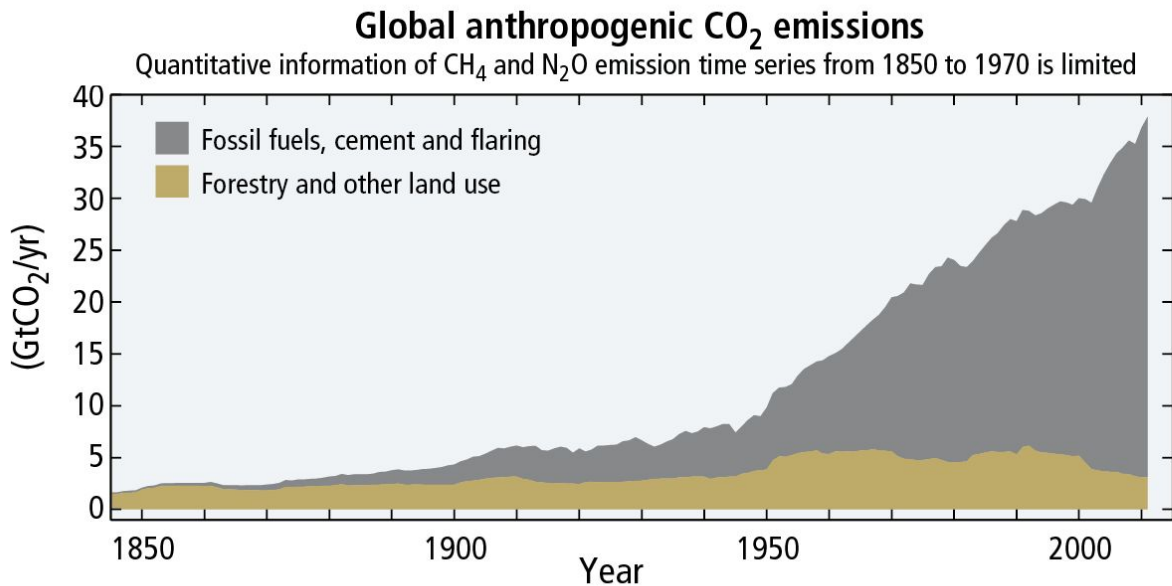
NWV

# Greenhouse gas concentrations in the atmosphere



Increase in greenhouse gases in the atmosphere are primarily from increases in carbon dioxide, methane and nitrous oxide.

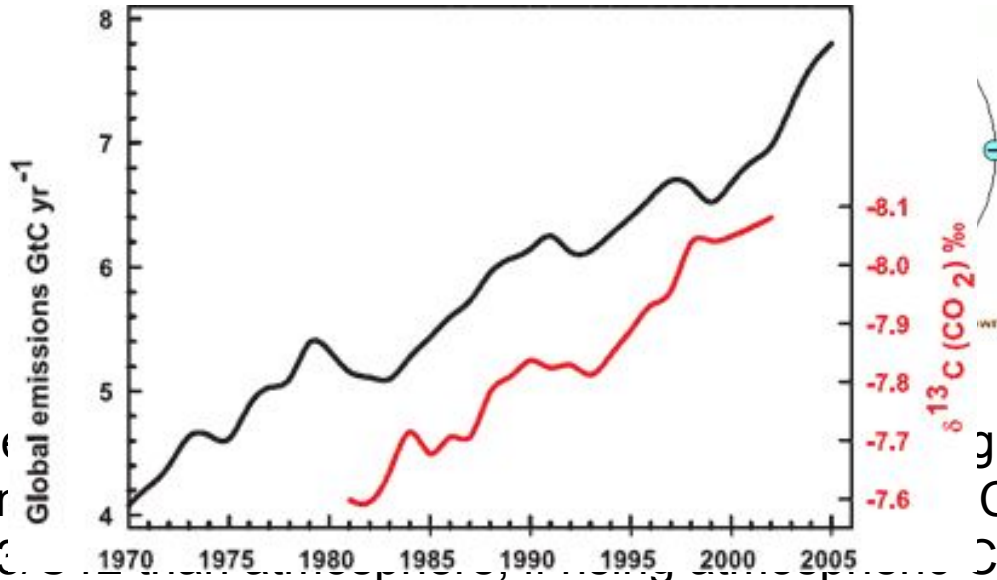
# Greenhouse gas changes in the atmosphere



Global anthropogenic CO<sub>2</sub> emissions

# Greenhouse gas changes in the atmosphere

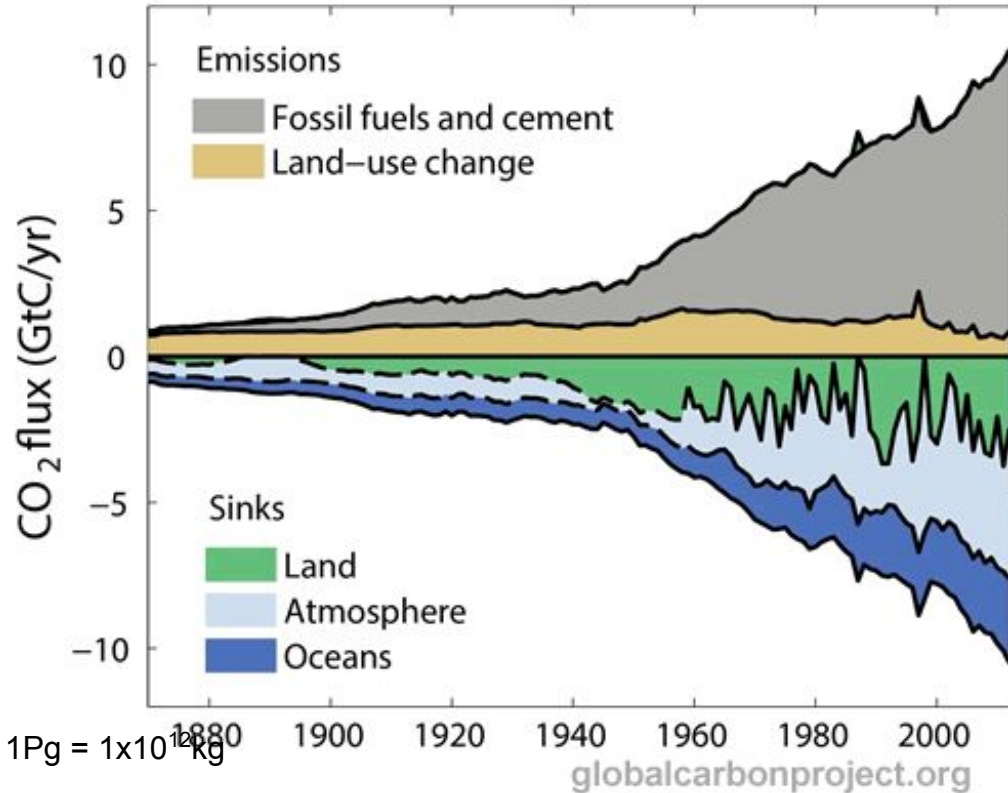
Can examine the ratio of different carbon isotopes, to show that CO<sub>2</sub> rise in atmosphere is from human fossil-fuel combustion.



Carbon is present in plants as C<sub>12</sub>, when plants are burned they have lower C<sub>13</sub>. From fossil fuels, C<sub>13</sub>/C<sub>12</sub> should be falling.

Lighter carbon C<sub>12</sub>. Plants CO<sub>2</sub> comes

# Carbon cycle

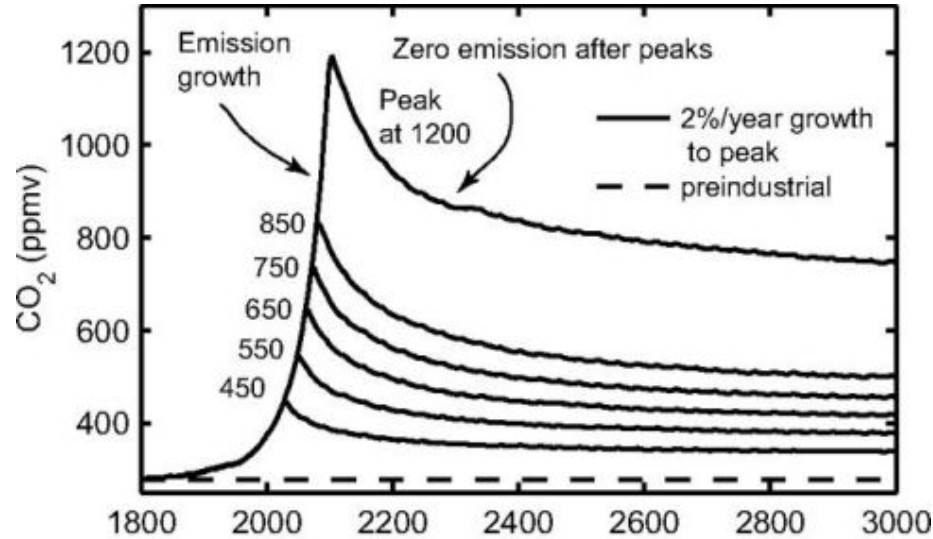


Global carbon budget from 1870 to 2012

Perturbation of the global carbon cycle caused by anthropogenic activities, averaged globally for the decade 2006–2015 (GtCO<sub>2</sub>/yr)



# Permanent CO<sub>2</sub> removal is slow



*Solomon et al., 2009, PNAS.*

Processes	Time scale (years)	Reactions
Land uptake: Photosynthesis–respiration	1–10 <sup>2</sup>	$6\text{CO}_2 + 6\text{H}_2\text{O} + \text{photons} \rightarrow \text{C}_6\text{H}_{12}\text{O}_6 + 6\text{O}_2$ $\text{C}_6\text{H}_{12}\text{O}_6 + 6\text{O}_2 \rightarrow 6\text{CO}_2 + 6\text{H}_2\text{O} + \text{heat}$
Ocean invasion: Seawater buffer	10–10 <sup>3</sup>	$\text{CO}_2 + \text{CO}_3^{2-} + \text{H}_2\text{O} \rightleftharpoons 2\text{HCO}_3^-$
Reaction with calcium carbonate	10 <sup>3</sup> –10 <sup>4</sup>	$\text{CO}_2 + \text{CaCO}_3 + \text{H}_2\text{O} \rightarrow \text{Ca}^{2+} + 2\text{HCO}_3^-$
Silicate weathering	10 <sup>4</sup> –10 <sup>6</sup>	$\text{CO}_2 + \text{CaSiO}_3 \rightarrow \text{CaCO}_3 + \text{SiO}_2$

# Summary

- Trace gases in the atmosphere ( $\text{H}_2\text{O}$ ,  $\text{CO}_2$ ,  $\text{CH}_4$ ,  $\text{N}_2\text{O}$ ...) are opaque to outgoing infrared radiation and responsible for the greenhouse effect.
- Due to greenhouse effect, the surface must warm enough to balance both incoming sunlight and the radiation from the atmosphere back to the surface
- Climate system has 'internal' and 'external' modes of variability that can span from 1-100,000 years.
- Increase in  $\text{CO}_2$  concentrations since from pre-industrial levels is from human fossil-fuel emissions, and permanent removal from the atmosphere is a slow process

# That's it for today!

Tomorrow in Climate 102 we will look at how to quantify the climate system response to both natural and human-caused forcings, and how scientists can detect and attribute observed changes in the climate system to human activity.

Other resources:

Check out

<http://globalchange.mit.edu/news-events/education#resources>

